

---

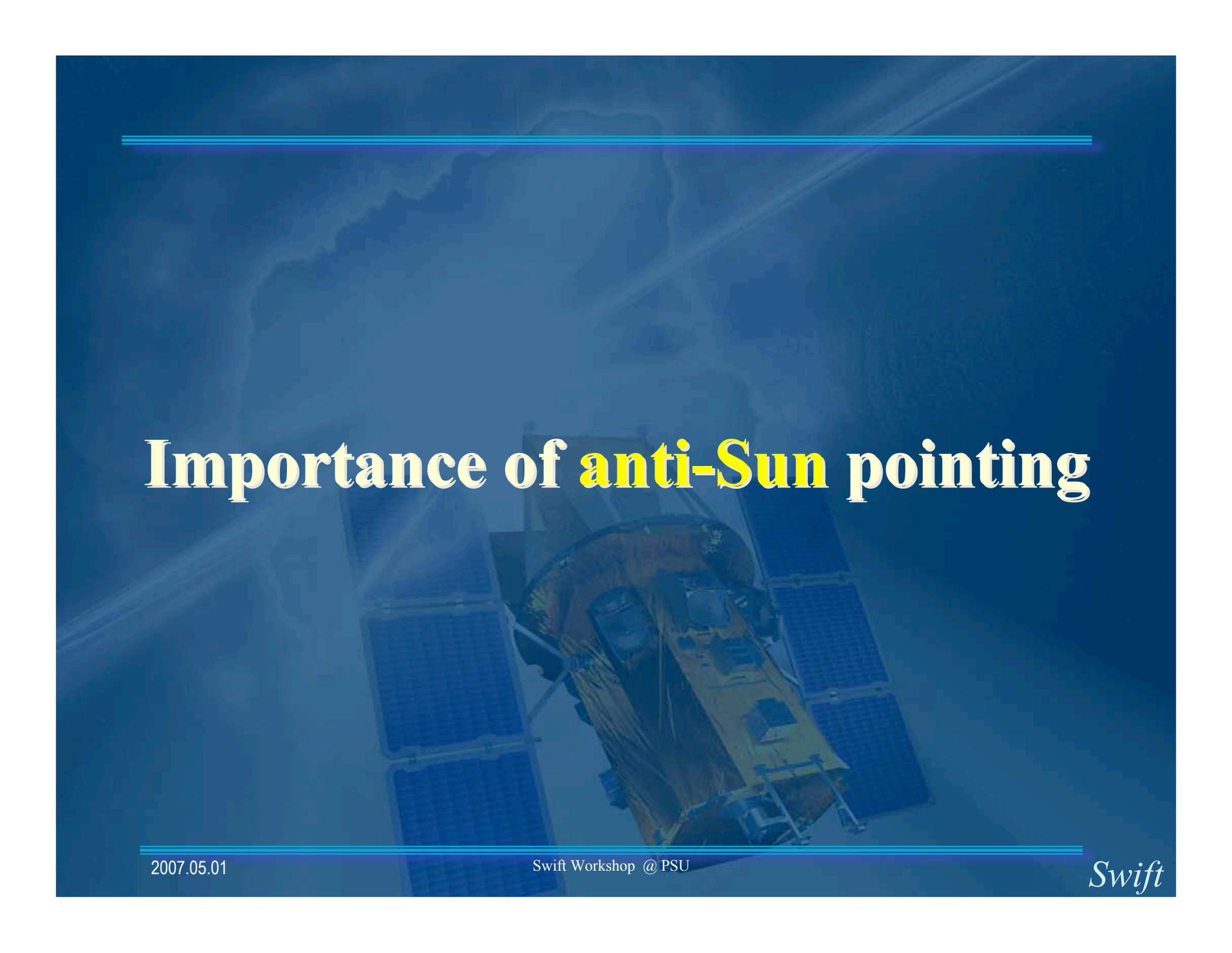
# Swift observation strategy: GRB Redshifts

T. Sakamoto (GSFC/ORAU)

# Contents

---

- Importance of **anti-Sun pointing**
- Screening “possible” high redshift GRBs from the BAT prompt emission properties



---

# Importance of **anti-Sun** pointing

# Importance of GRB redshifts

- **High-z GRBs**      Huge potential for studying early universe ( $z > 10$ )

1. IR detection and spectroscopy
2. Deep optical observations
3. Notify a “possible high-z” from the Swift data

- **Luminosity indicators**      Measurement of Dark energy (Schaefer)

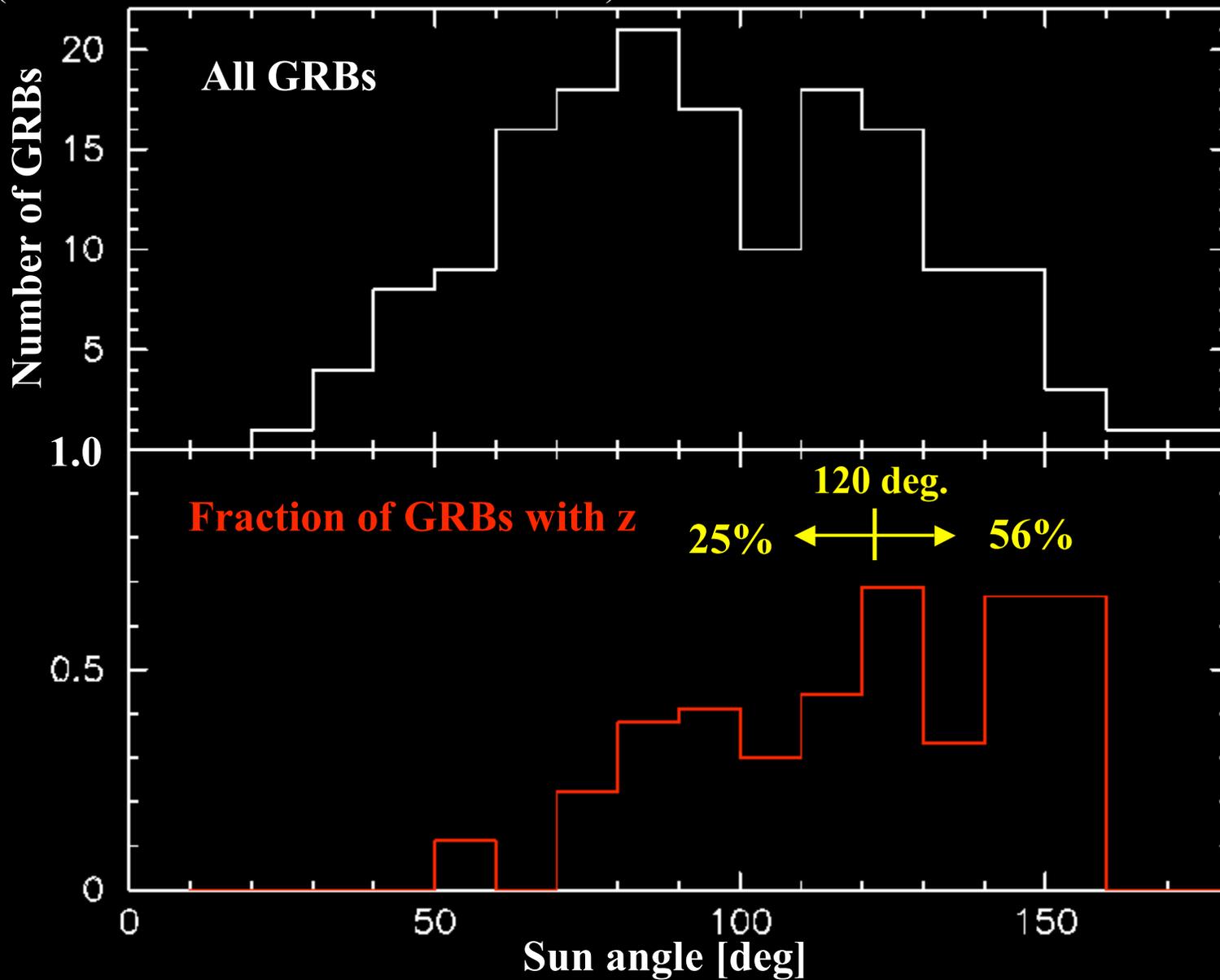
GRB properties ( $E_p$ ,  $V$ ,  $Lag$ , ..) + redshift

Variety of redshifts (from low to high  $z$ )

GRBs at the **anti-Sun** directions

# Sun angle vs. Redshift

(from GRB 050502B to GRB 061222B)



# Swift observation strategy

---

---

- XRT positions
- Bursts with anti-Sun directions

- **Point Swift to anti-Sun position as far as we can.**
- **XRT will focus on the prompt localization + early X-ray AG.**
- **No long follow-up by Swift (Sun angle  $< 120$  deg. bursts)**

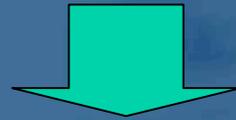


---

# Screening “possible” high redshift GRBs from the BAT prompt emission properties

# Screening “possible” high-z GRBs from the BAT data

Find common characteristics of high-z GRBs

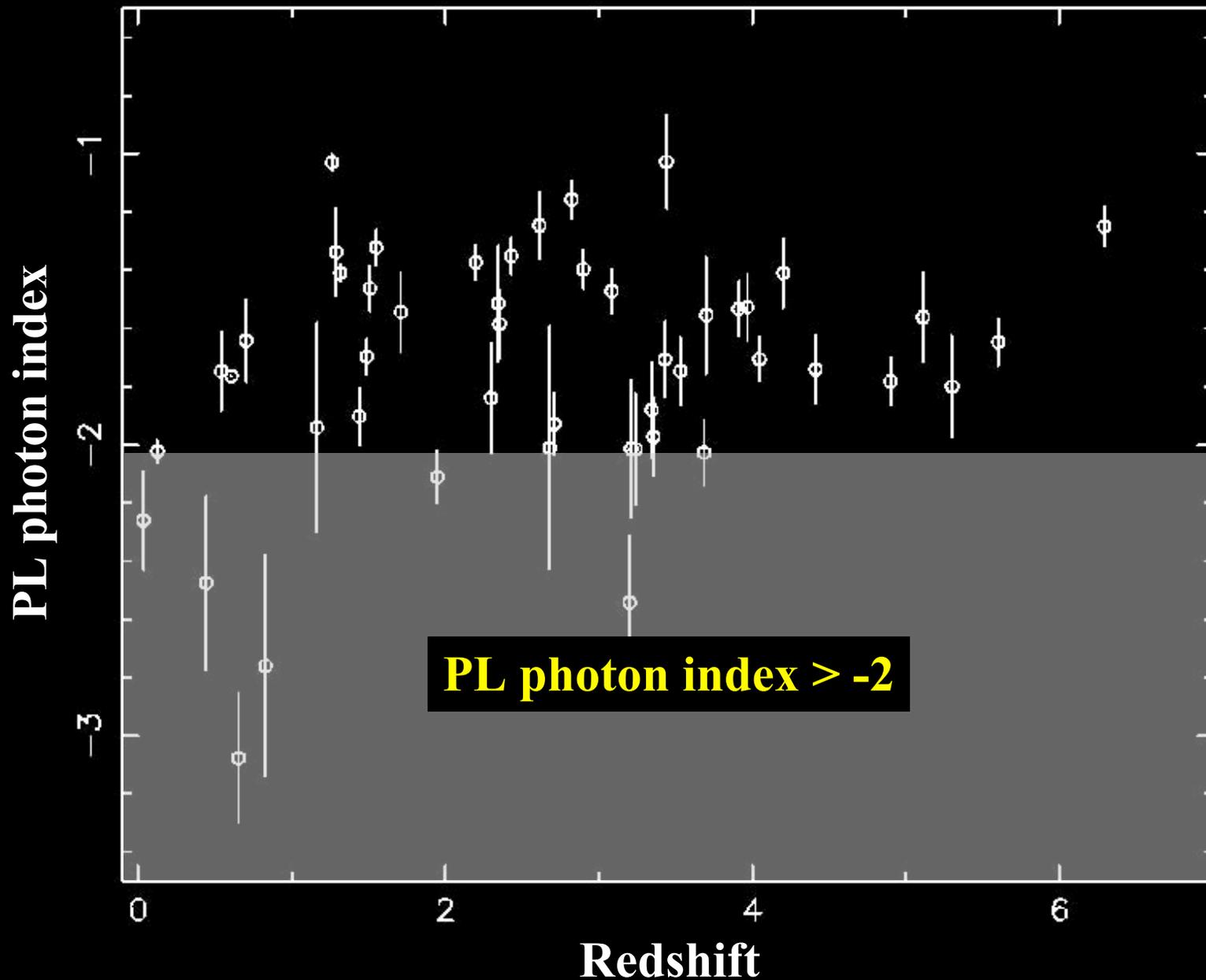


Alert the ground observers as a “possible” high-z GRB

## Sample of GRBs

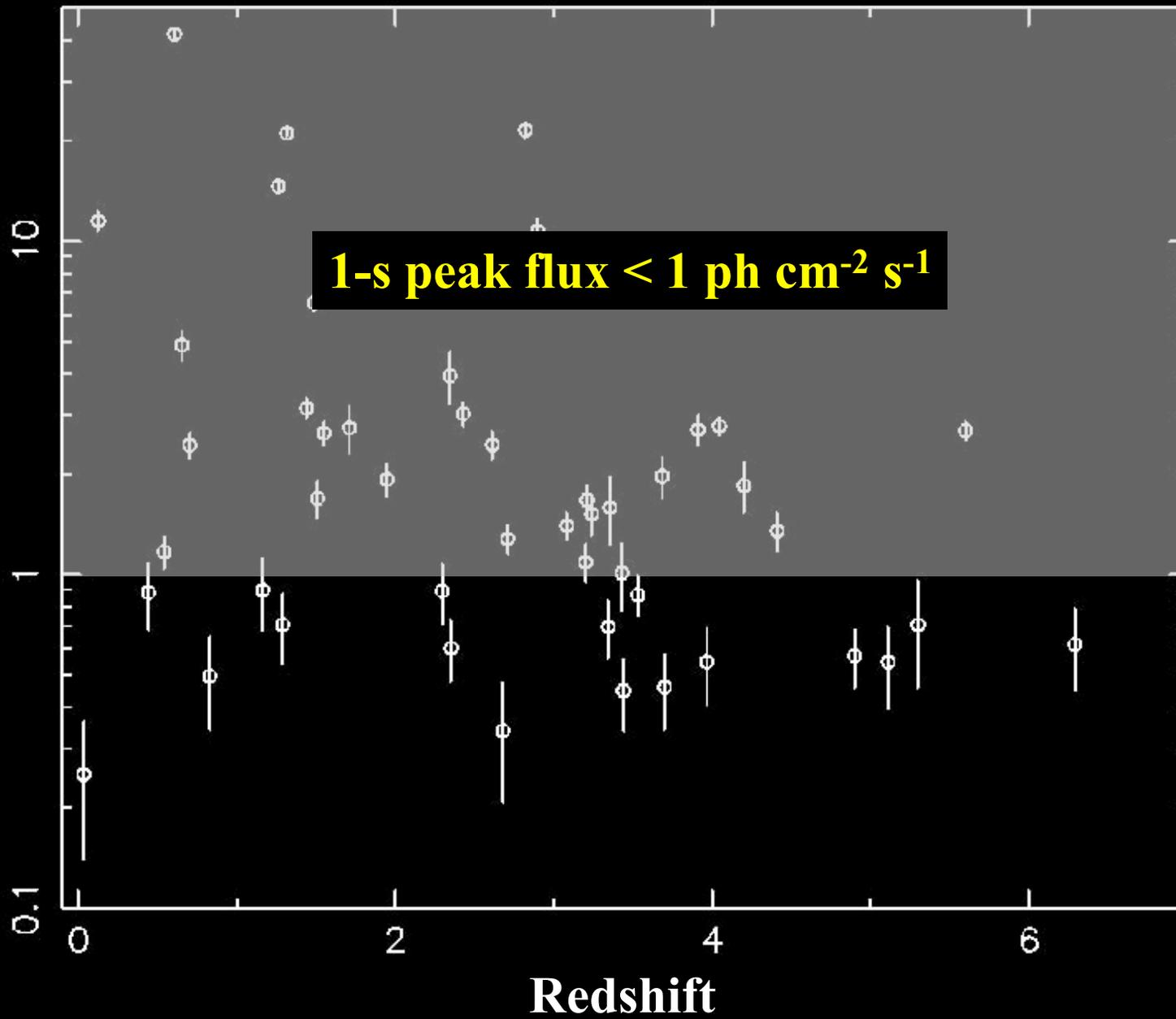
- 49 known-z long GRBs ( $T_{90} > 2$  sec)
- 188 long GRBs (Dec 2004 – Feb 2007)
- BAT event data analysis (*batgrbproduct*)

# Redshift vs. PL photon index

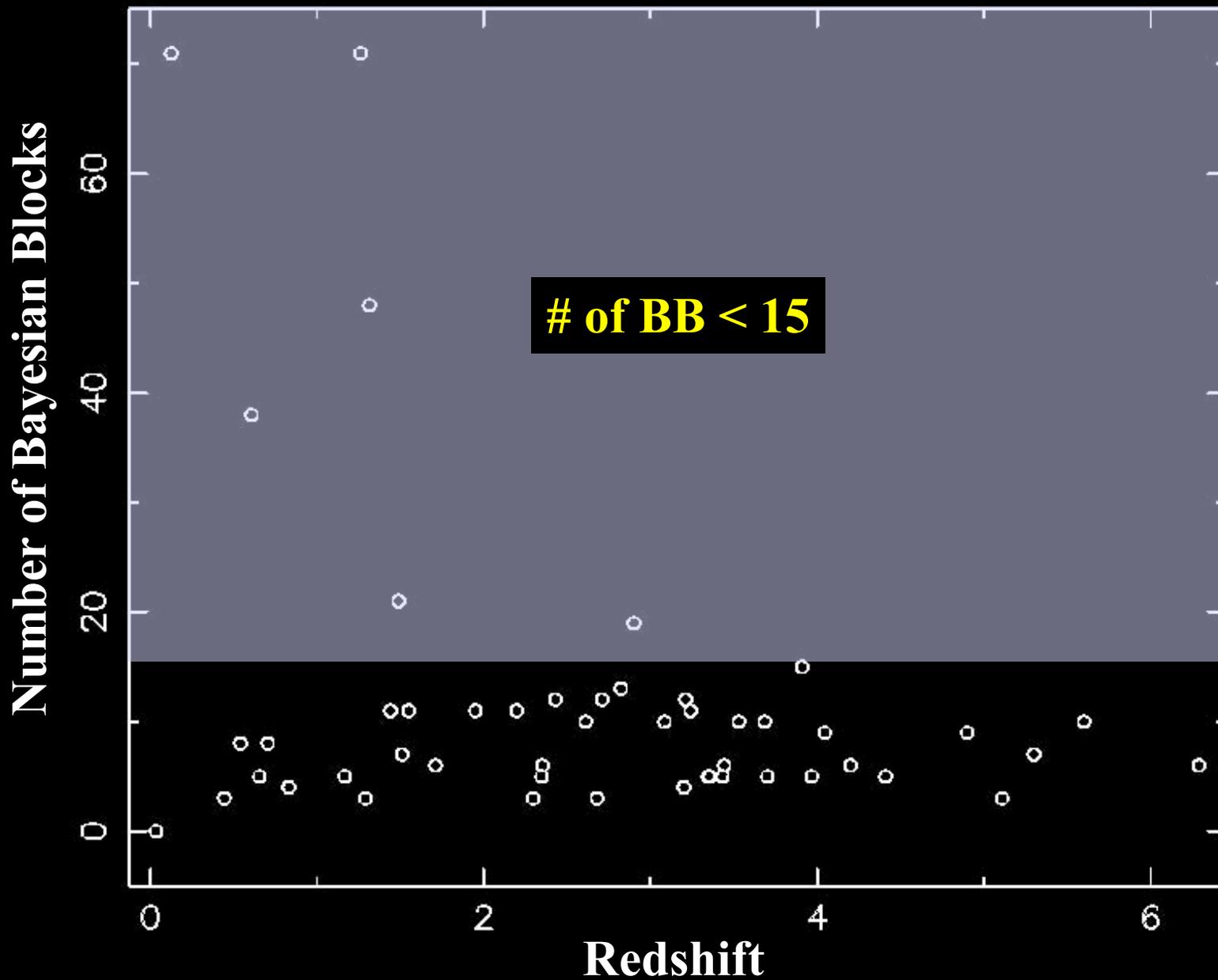


# Redshift vs. 1-s peak photon flux

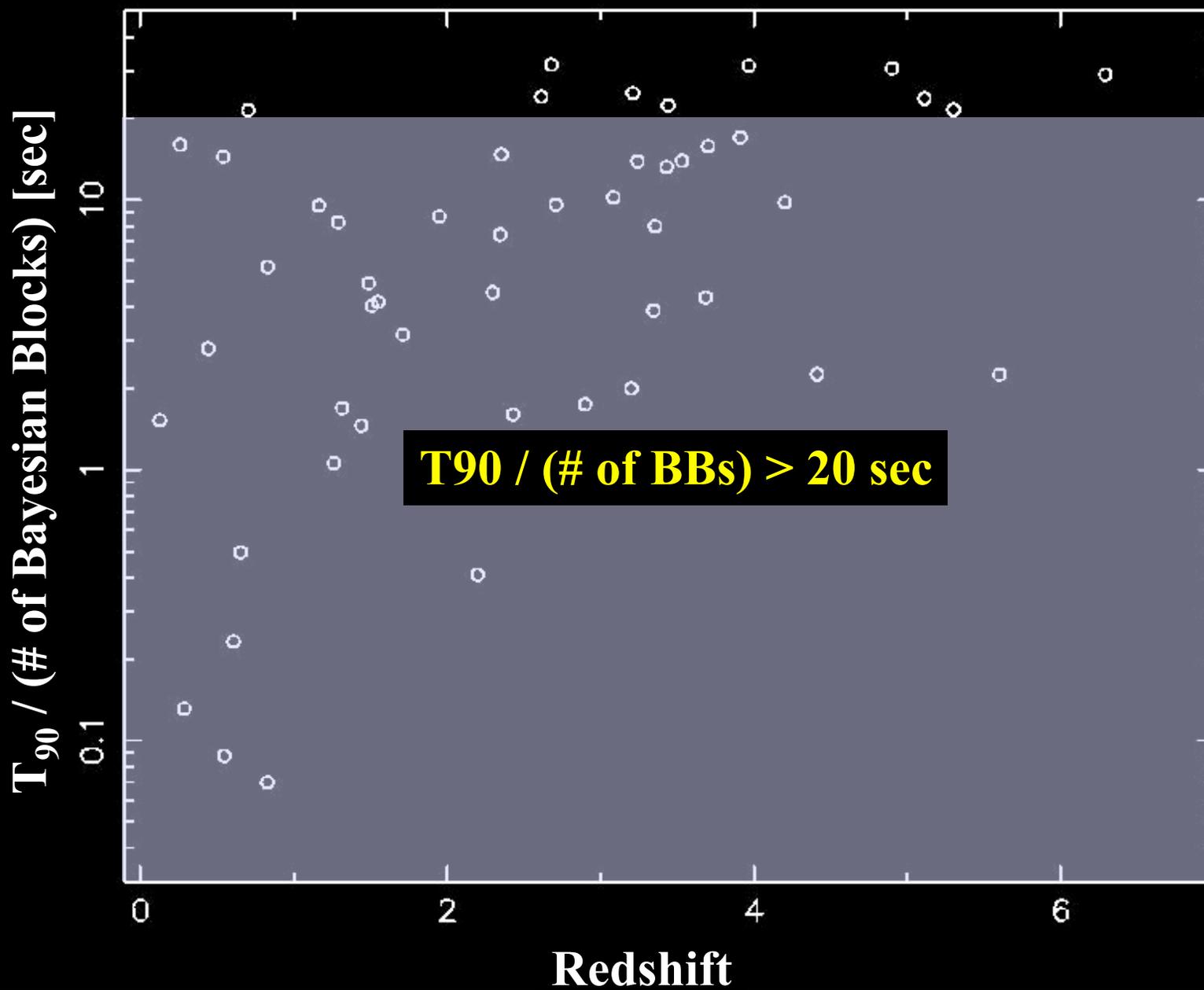
1-s peak photon flux (15-150 keV) [ $\text{ph cm}^{-2} \text{s}^{-1}$ ]



# Redshift vs. # of Bayesian Blocks



# Redshift vs. $T_{90} / (\# \text{ of BBs})$



# Screening high-z GRBs

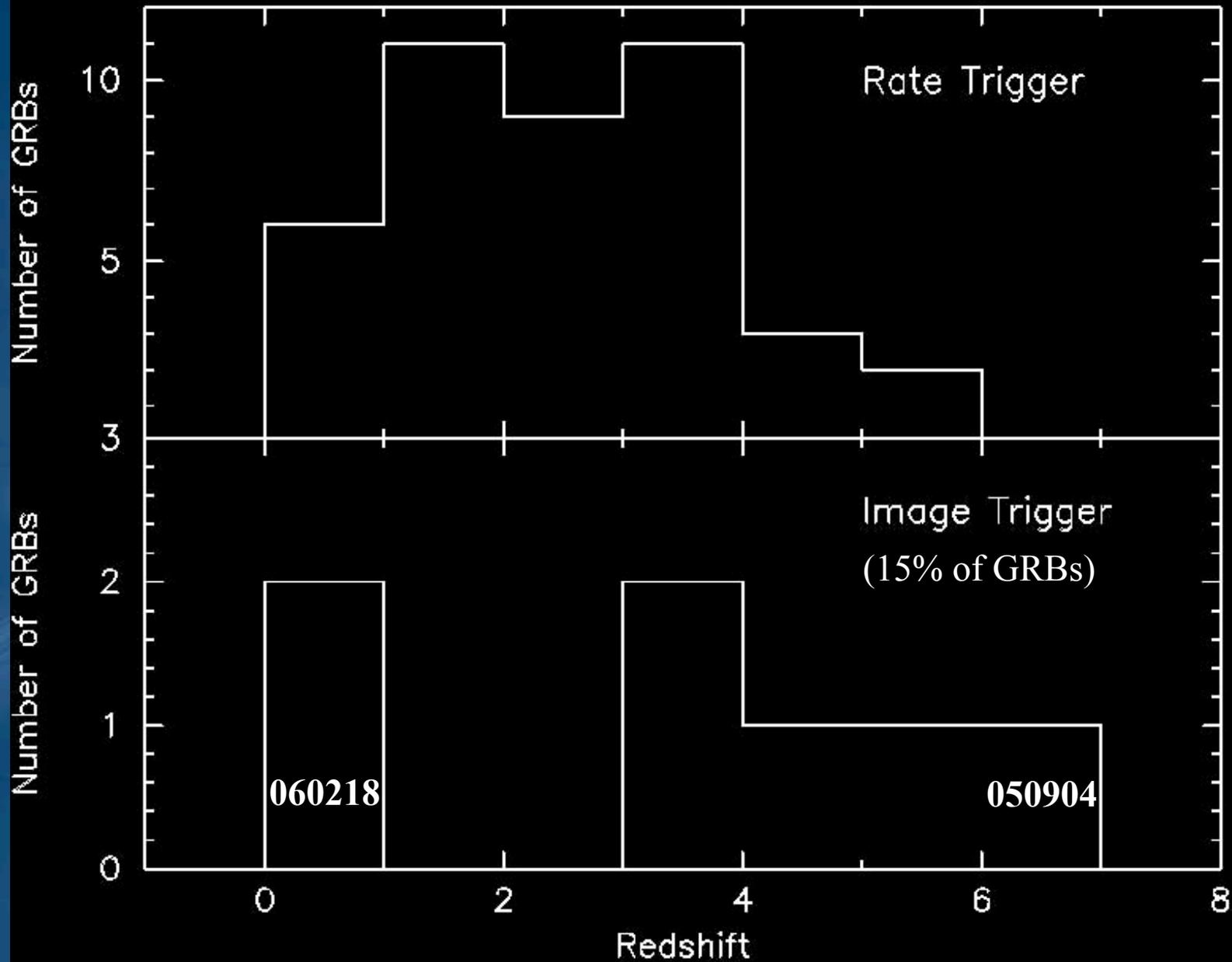
- PL photon index  $> -2$
- 1-s peak flux  $< 1 \text{ ph cm}^{-2} \text{ s}^{-1}$
- # of BBs  $< 15$
- $T90 / (\text{\# of BBs}) > 20 \text{ sec}$

GRB	z	PhIndex	Peak flux	BB	T90/BB
050730	3.967	-1.53	0.5	5	31.3
050814	5.3	-1.80	0.7	7	21.6
050904	6.29	-1.25	0.6	6	29.0
060510B	4.9	-1.78	0.6	9	30.6
060522	5.11	-1.56	0.5	3	23.7
061110B	3.44	-1.03	0.4	6	22.3

[Excluded GRB with  $z > 4.5$  is 060927 ( $z = 5.6$ )]

**6 GRBs / 49 known-z long GRBs (11 GRBs / 188 long GRBs)**

# BAT image trigger



# Summary

## ■ Point Swift to the **anti-Sun direction**

### **Double the number of redshifts**

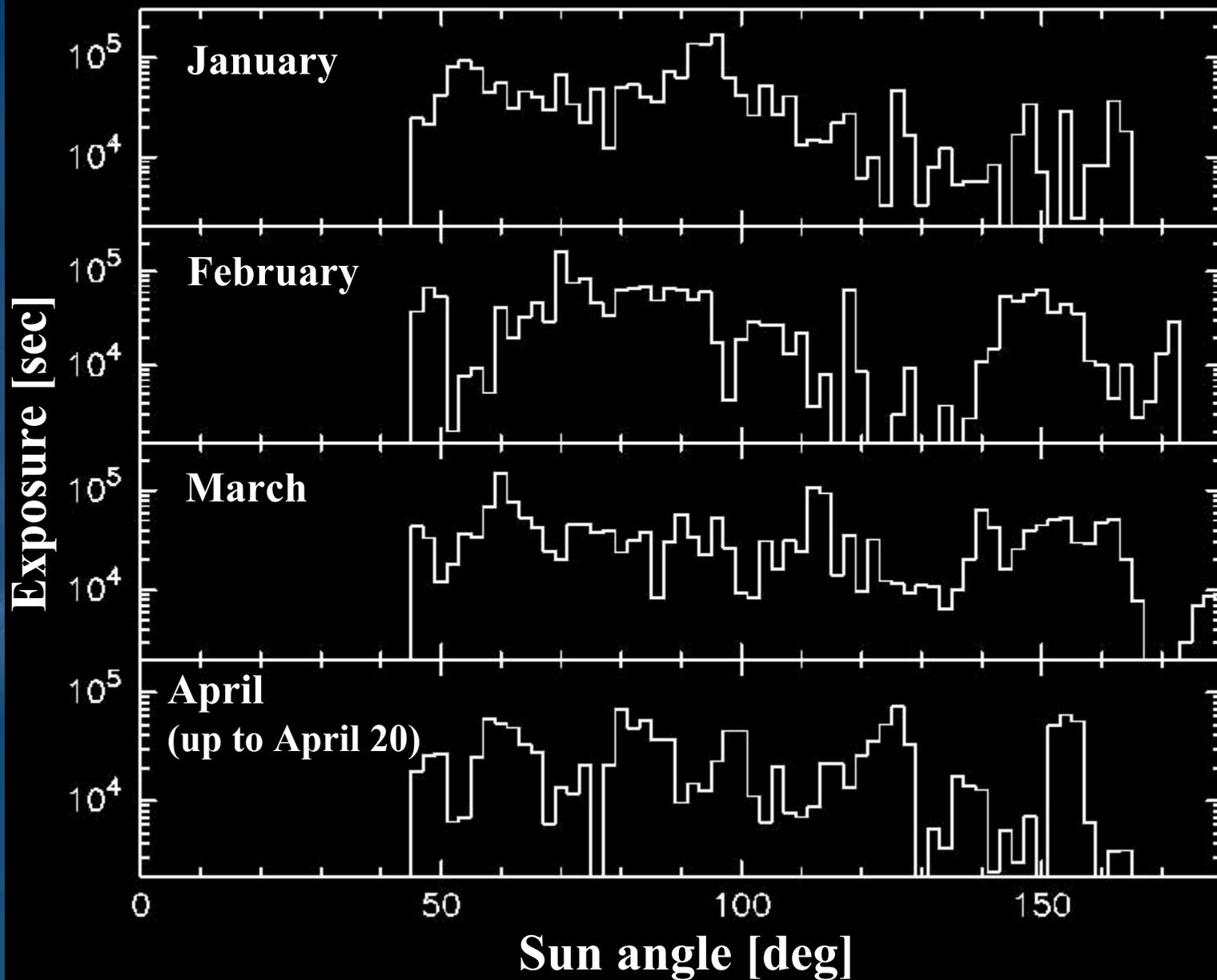
- XRT: prompt position but no long follow-up
- Q1. How long do we need to cool XRT to get the position?  
(Is cooling XRT during the SAA period enough?)
- Q2. Is it possible to change the threshold of a mode switching?

## ■ Screening “possible” high redshift GRBs from the BAT prompt emission properties

- We can **alert possible high- $z$  ( $z > 4$ ) GRBs** ( $\sim 5$  GRBs year<sup>-1</sup>)
- GRB detected by the **BAT image trigger** is very likely to be an “interesting” burst



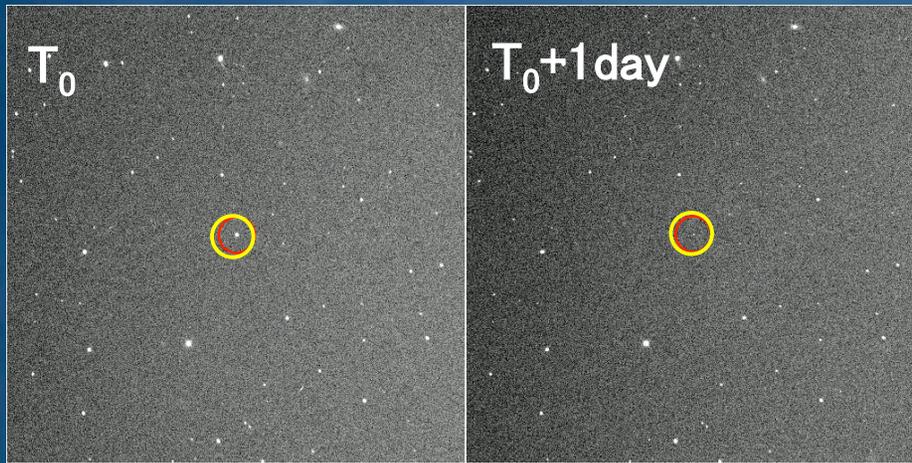
# Sun angle distribution 2007



# City lights are not an excuse

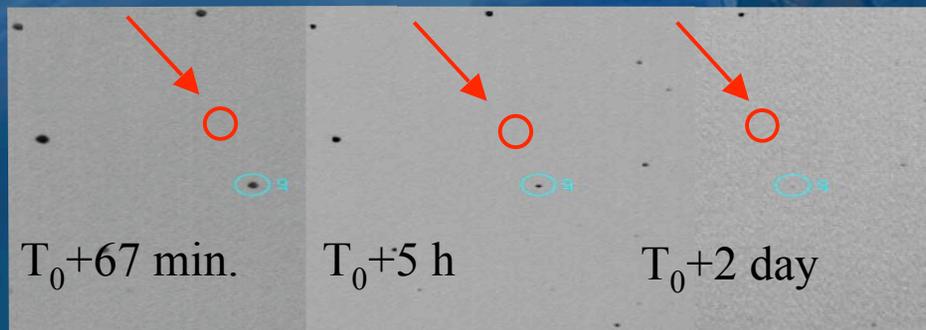
## Optical observations of GRB 030329

Roof of RIKEN main building (Wako, Saitama)



(Torii)

Roof of Tokyo Tech. main building (Meguro, Tokyo)

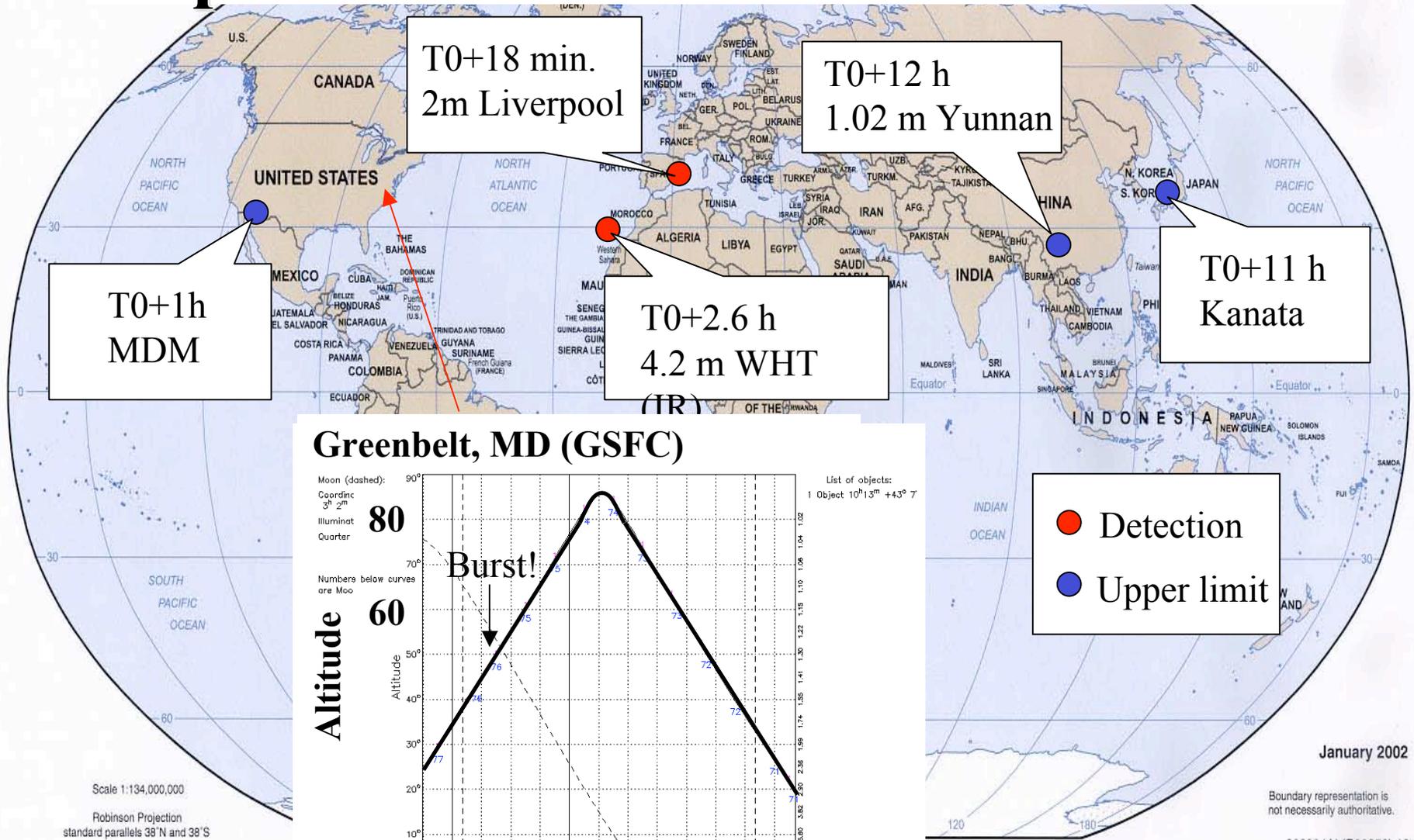


(Sato et al.)

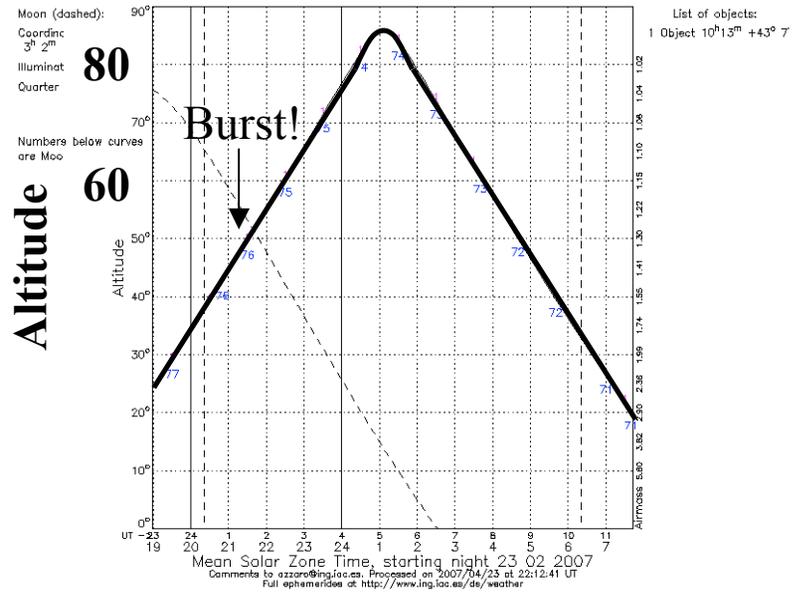
## Astronomy Picture of the Day (11/27/00)



# GRB 070223: perfect GRB from U.S. east coast



## Greenbelt, MD (GSFC)

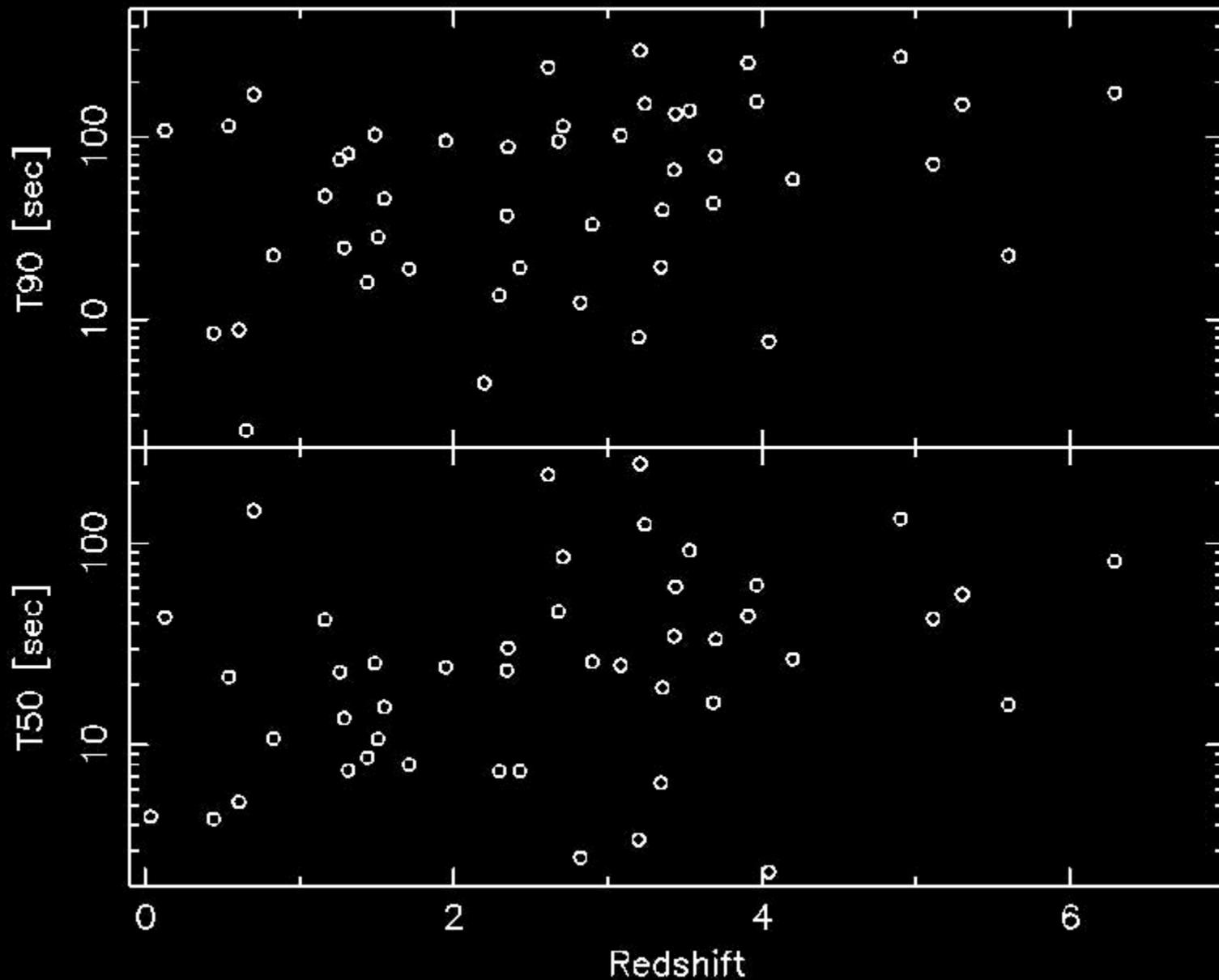


January 2002

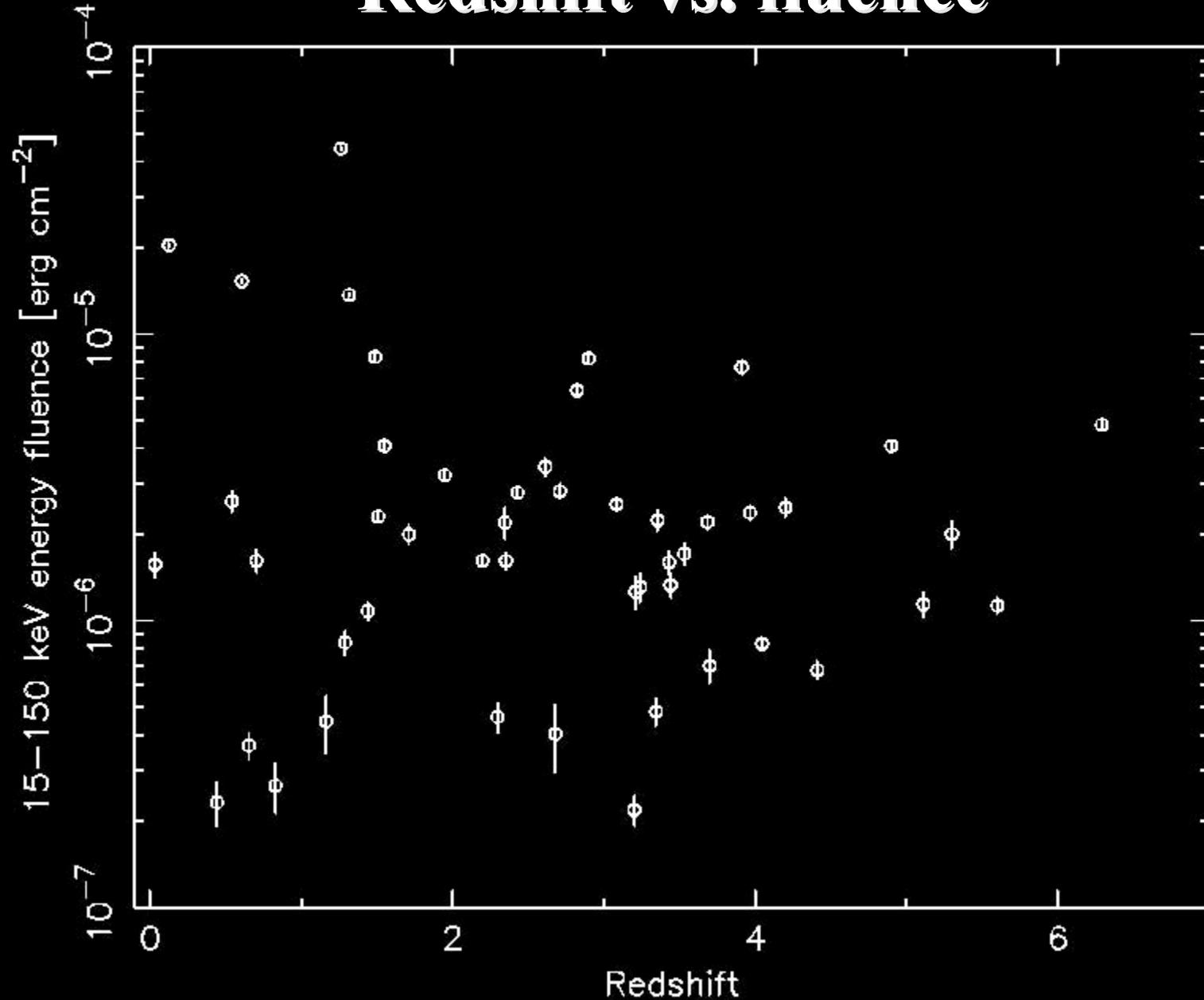
Boundary representation is not necessarily authoritative.

802804AI (R00352) 12-01

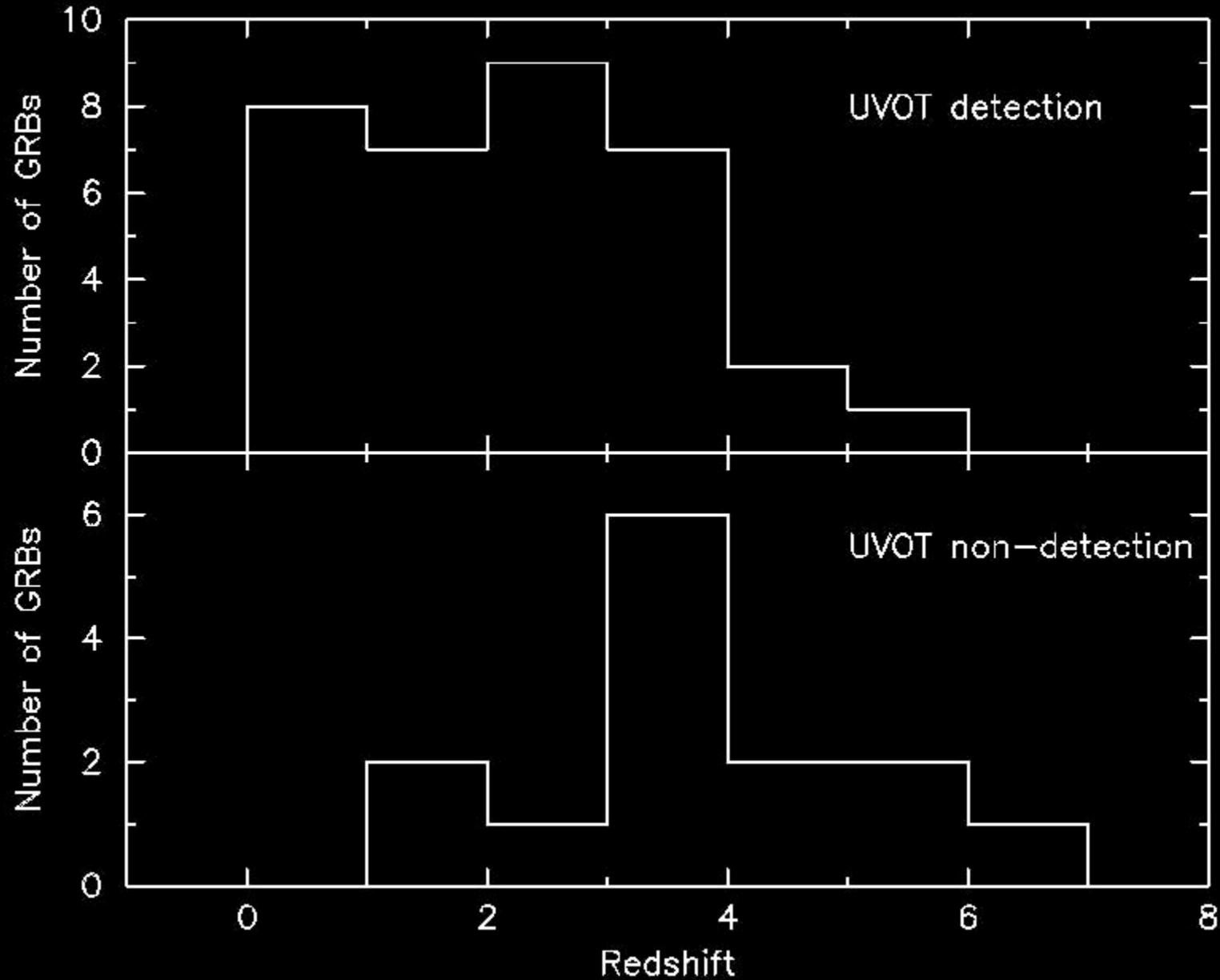
# Redshift vs. T90/T50



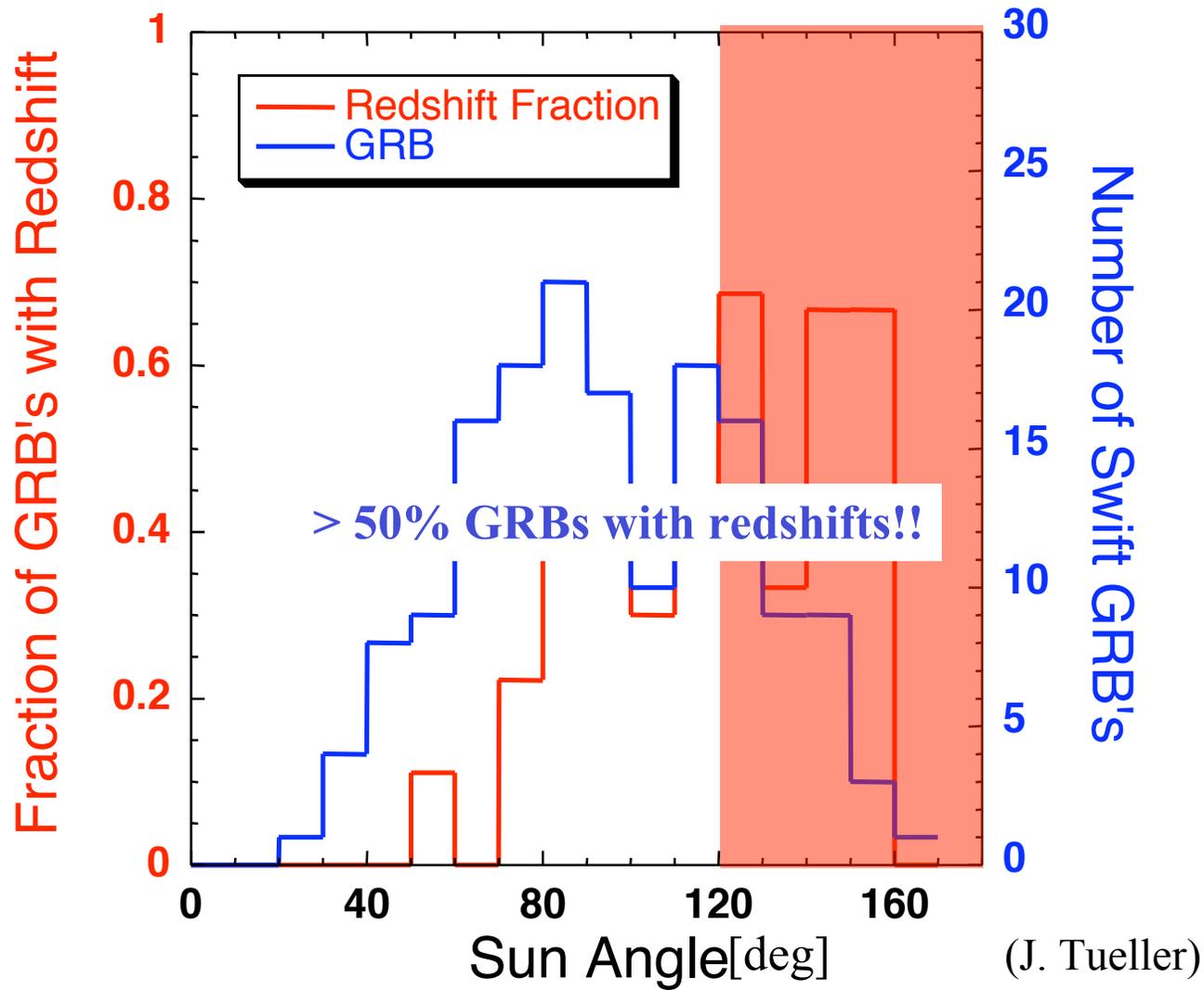
# Redshift vs. fluence



# Redshift vs. UVOT AG detection



# Sun angle vs. redshifts



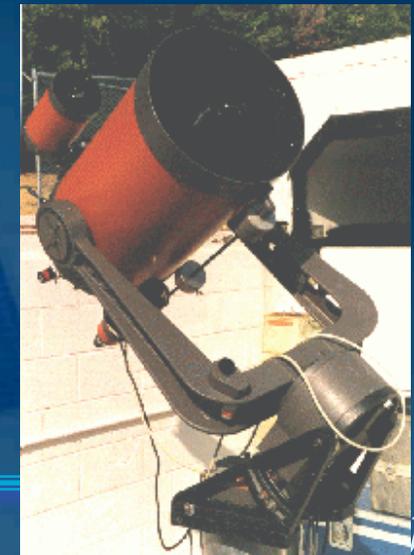
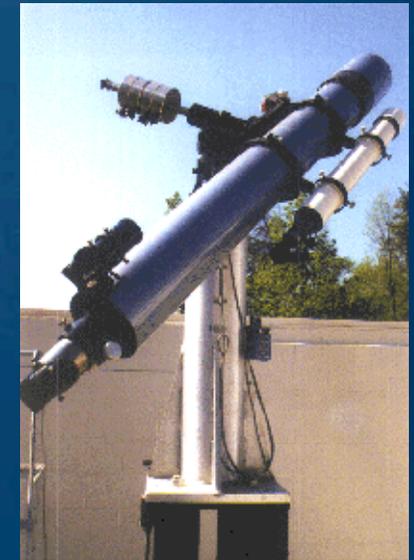
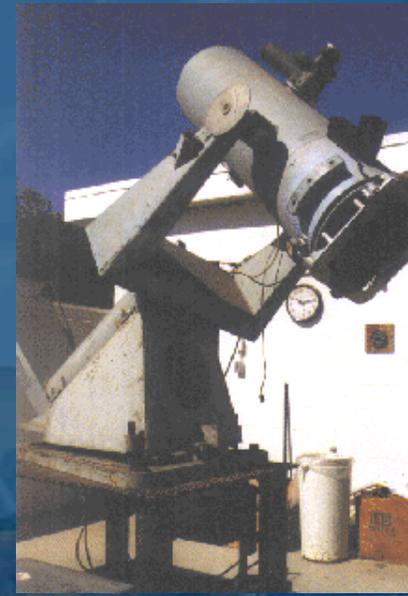
# Telescopes around GSFC

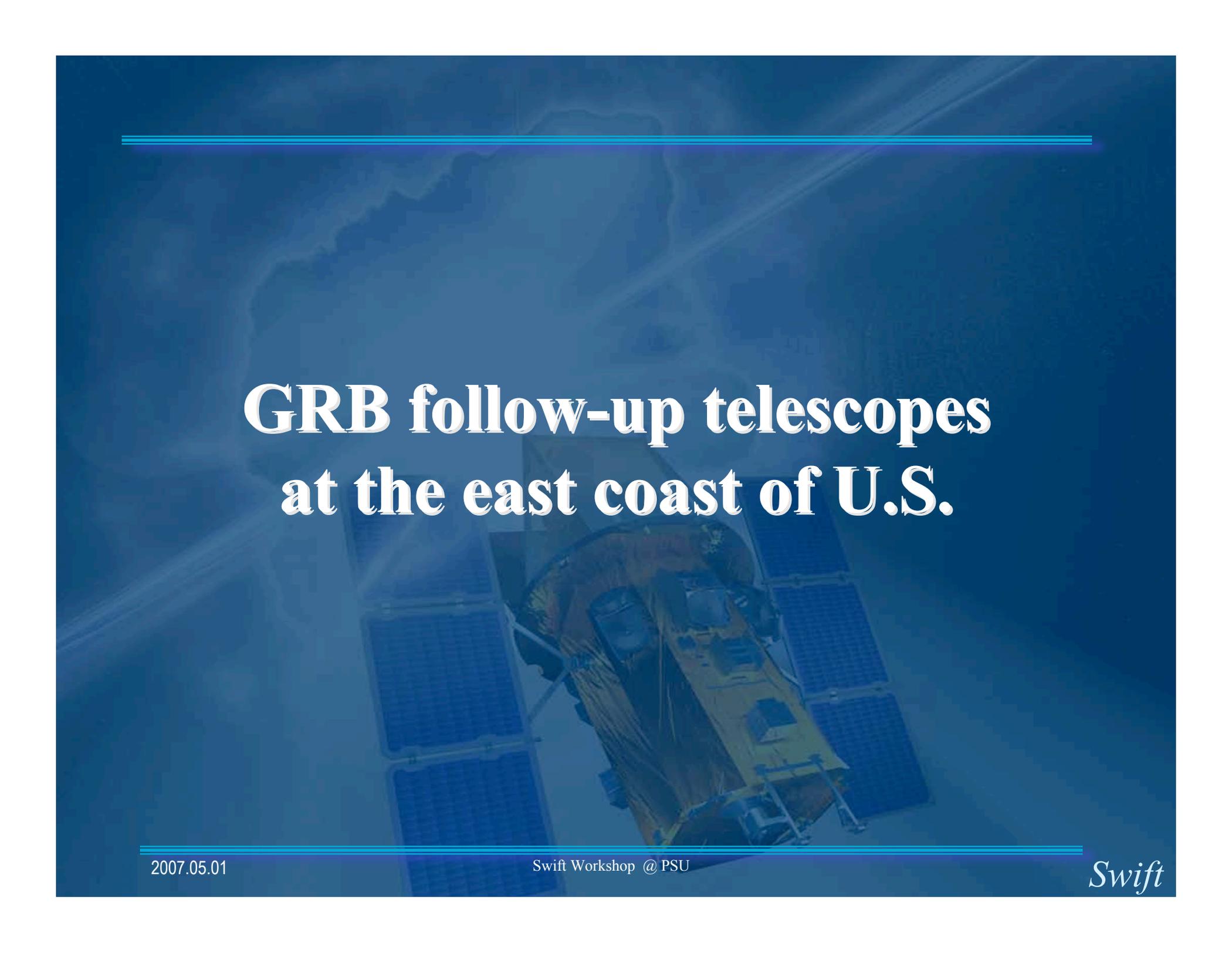
UMBC/Joint Center for Astrophysics

University of Maryland Observatory



0.83 m (top of physics building)

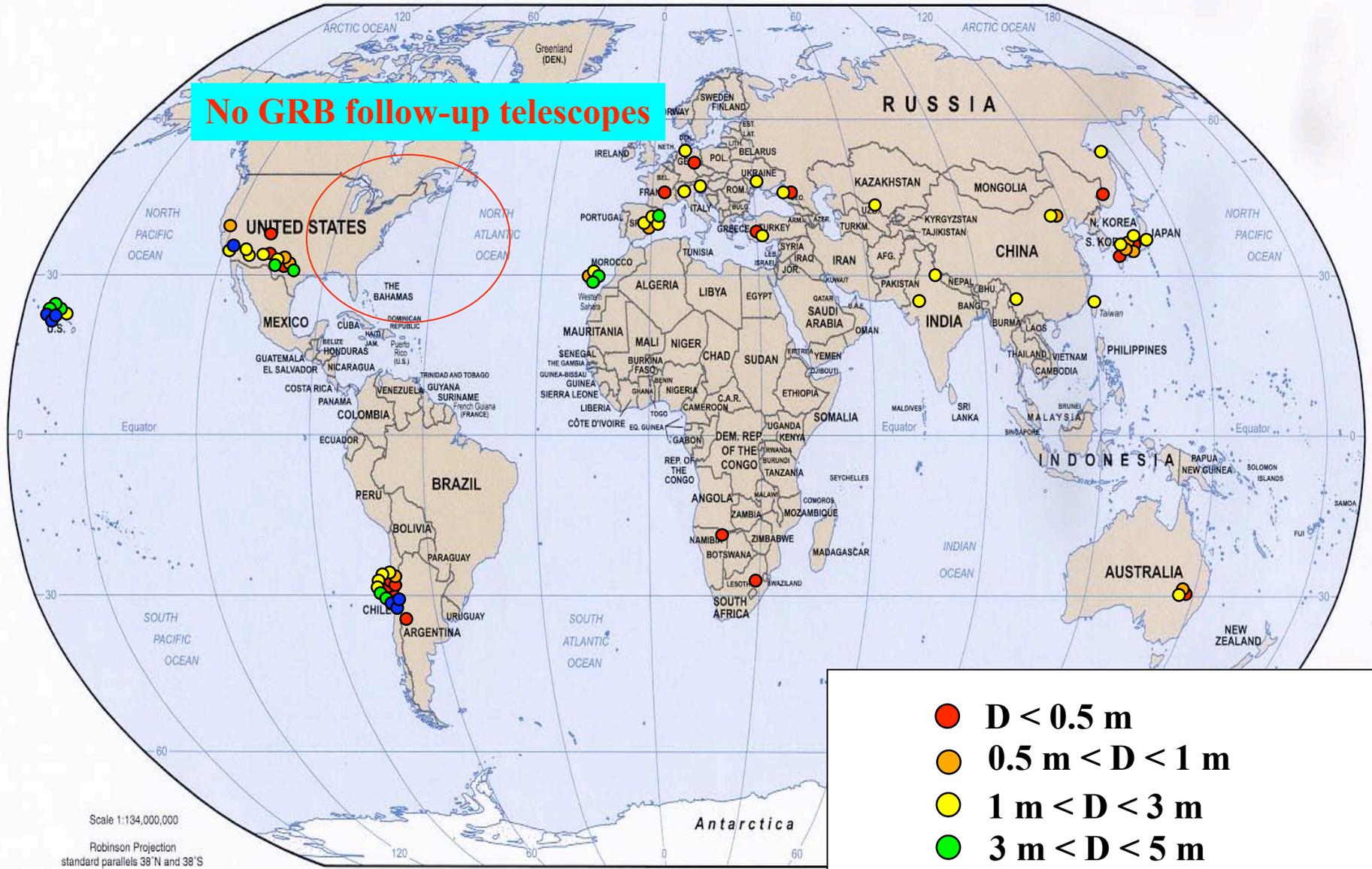




---

# GRB follow-up telescopes at the east coast of U.S.

# GRB follow-up telescopes

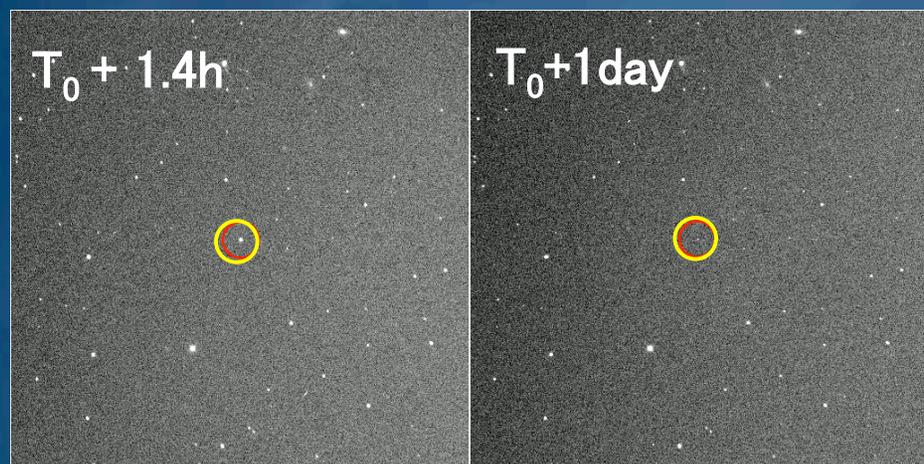


(From GCN Circular: September 2006 – March 2007)

# City lights are not an excuse

## Optical observations of GRB 030329

Roof of RIKEN main building (Wako, Saitama)

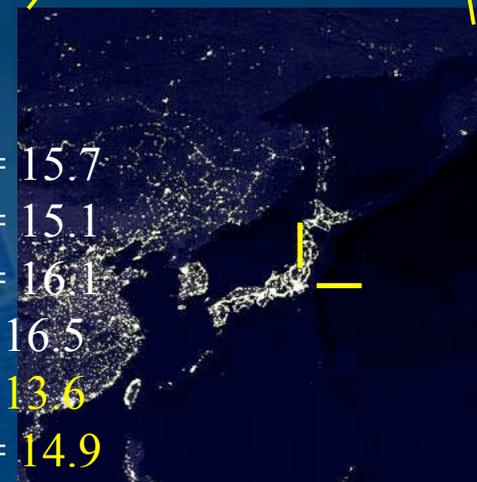


(0.25m Torii)

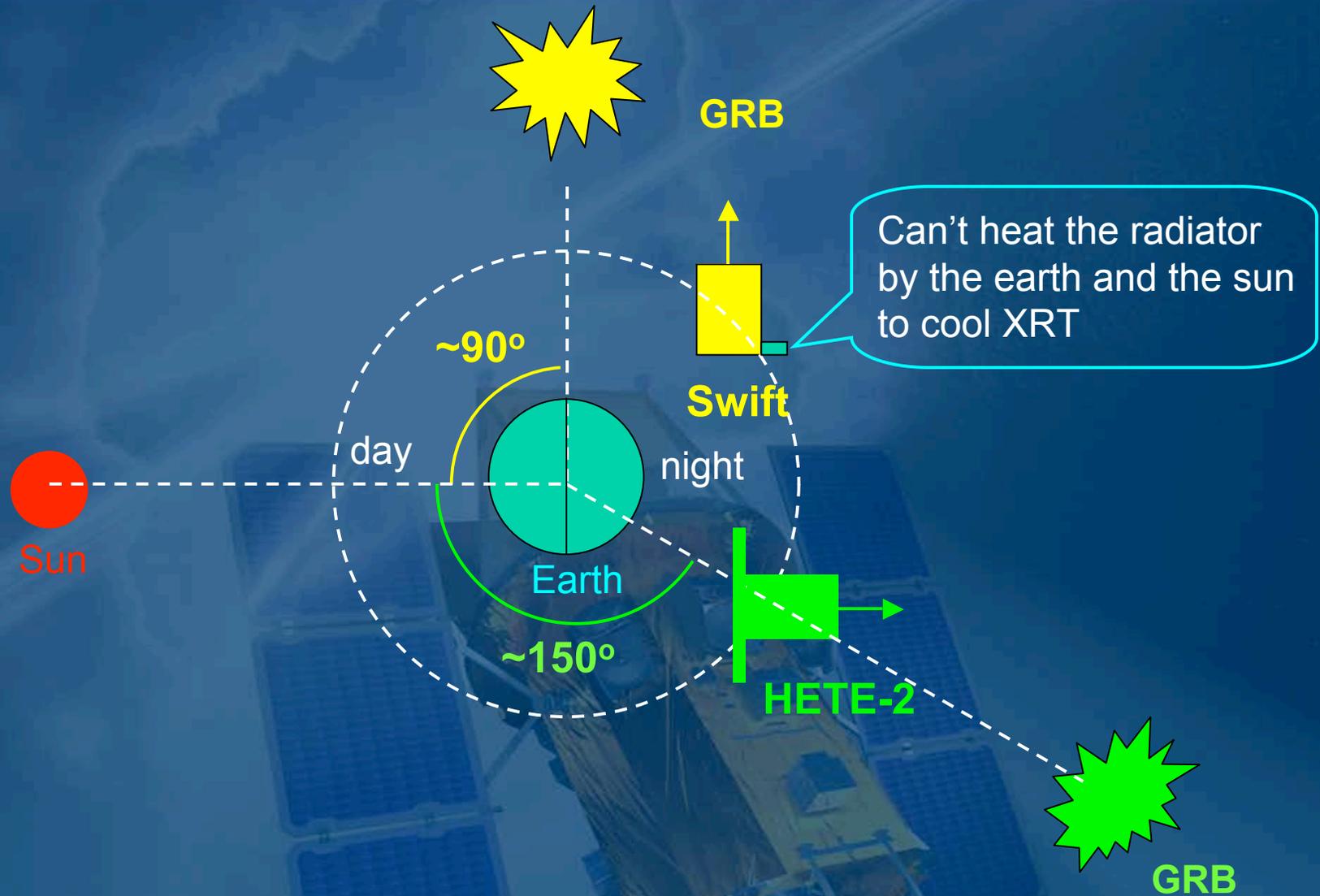
Swift GRBs (magnitude < 17 mag)

- |                          |                          |                          |
|--------------------------|--------------------------|--------------------------|
| - 050319A : unflt = 16   | - 051109A : unflt = 15.4 | - 060607A : white = 15.7 |
| - 050525A : unflt = 14.7 | - 051111A : unflt = 13   | - 060908A : white = 15.1 |
| - 050603A : R = 16.5     | - 051211B : unflt = 16.2 | - 060912A : white = 16.1 |
| - 050801A : unflt = 15   | - 060117A : R = 11.5     | - 060927A : unflt = 16.5 |
| - 050820A : R = 14.7     | - 060418A : V = 14.5     | - 061007A : unflt = 13.6 |
| - 050822C : V = 15.5     | - 060512A : white = 16.2 | - 061021A : white = 14.9 |
| - 050922C : unflt = 14.7 | - 060605A : unflt = 16.3 | - 070318A : V = 15.4     |

## Astronomy Picture of the Day (11/27/00)



# Why Swift is pointing closer to the Sun

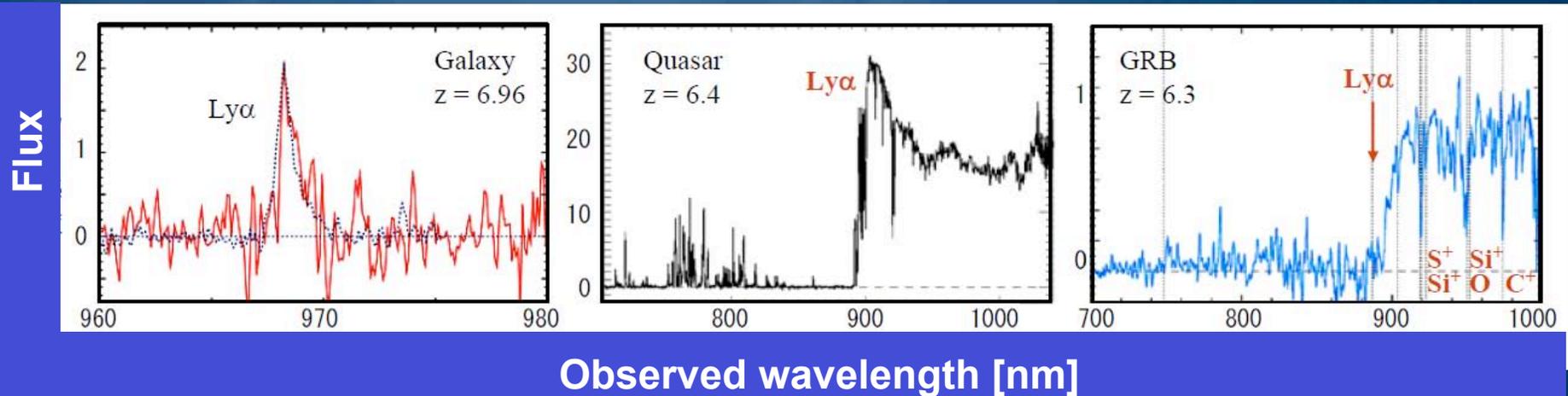


# High-z objects: Galaxy, Quasar, ... and GRB

Galaxy (Ly $\alpha$  Emitter)

Quasar

GRB



(Ohta Kyoto GRB workshop)

Galaxy:

- Very dim (low SN)
- Limited spectral information

Quasar:

- Not a 'normal' environment
- Ionized by own UV radiations
- Complicated spectrum

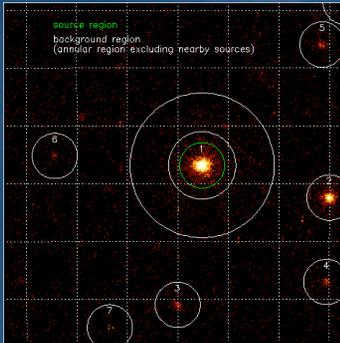
GRB:

- 'Normal' (star forming) environment
- Transient emission
- Simple power-law spectrum

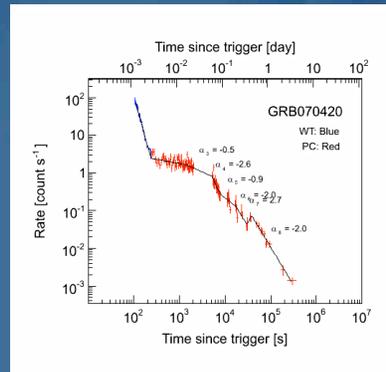
# Swift GRB observation strategy

## XRT position

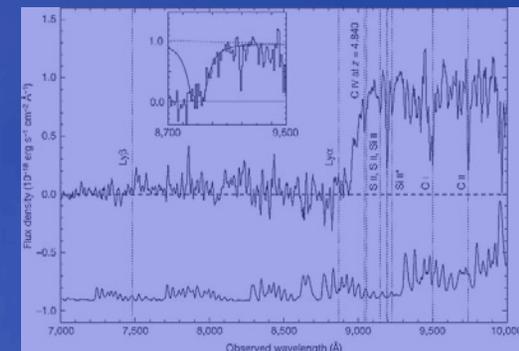
A



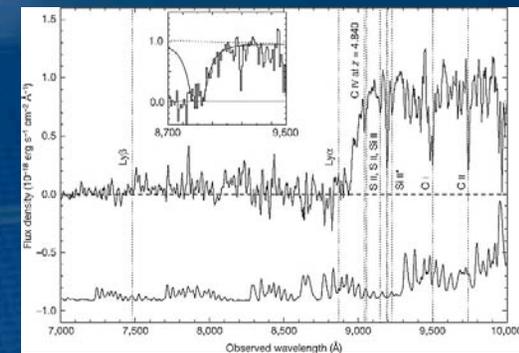
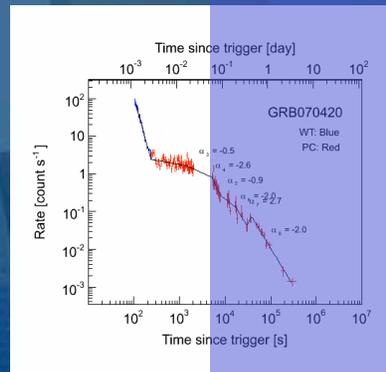
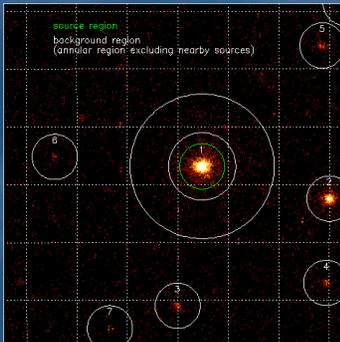
## XRT light curve



## Spectroscopy



B



- Point Swift to anti-Sun position as far as we can.
- XRT will focus on the prompt localization + early X-ray AG.
- No long follow-up by XRT (Sun angle < 120 deg. bursts)

# Advantage of GRBs for measuring dark energy

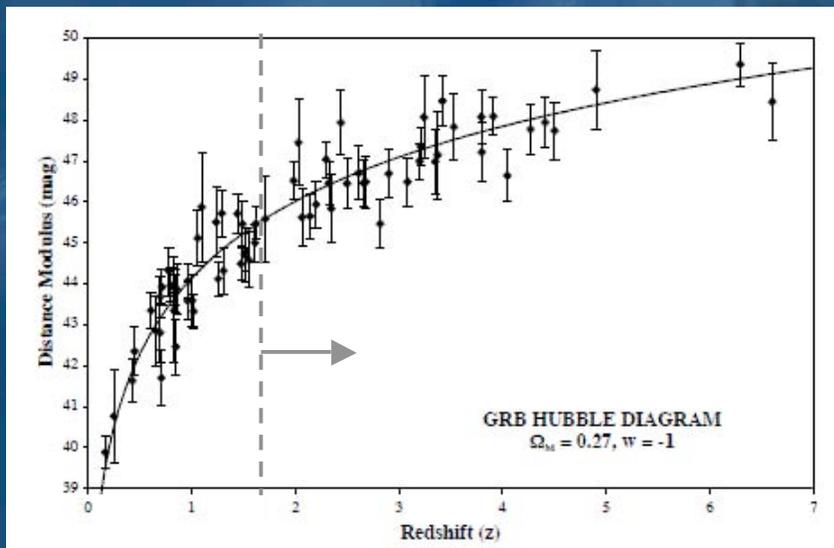
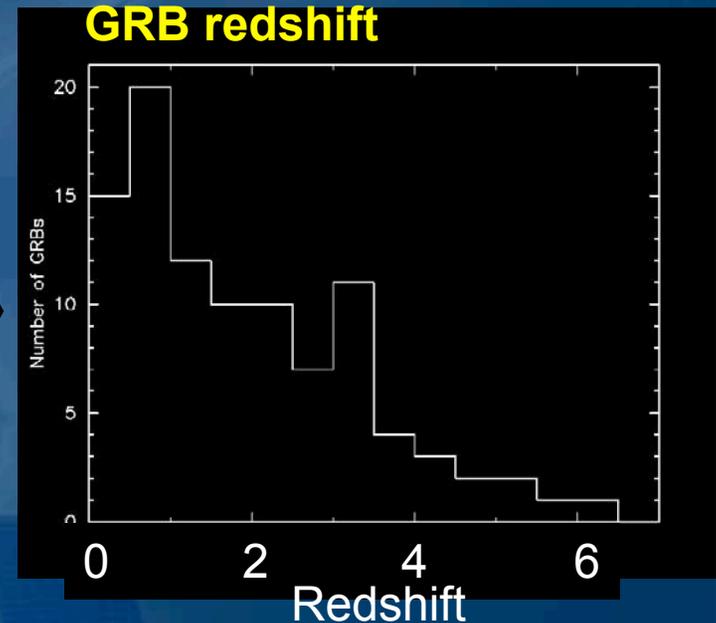
## GRB

Luminosity indicators + Redshift

- $E_{\text{peak}} - E_{\text{iso}}$
- $E_{\text{peak}} - L_{\text{iso}}$
- $E_{\text{peak}} - E_{\text{gamma}}$
- $\text{Lag} - L_{\text{iso}}$
- Variability -  $L_{\text{iso}}$
- etc...



Similar method to SN Ia with a **great advantage**



However...

**GRB luminosity indicators are premature.**

We need to..

- **Confirm the GRB luminosity indicators**
- **Calibrate the indicators**

**GRB can be used to measure dark energy!**

# Importance of GRB redshifts

**- High-z GRBs**      Huge potential for studying early universe ( $z > 10$ )

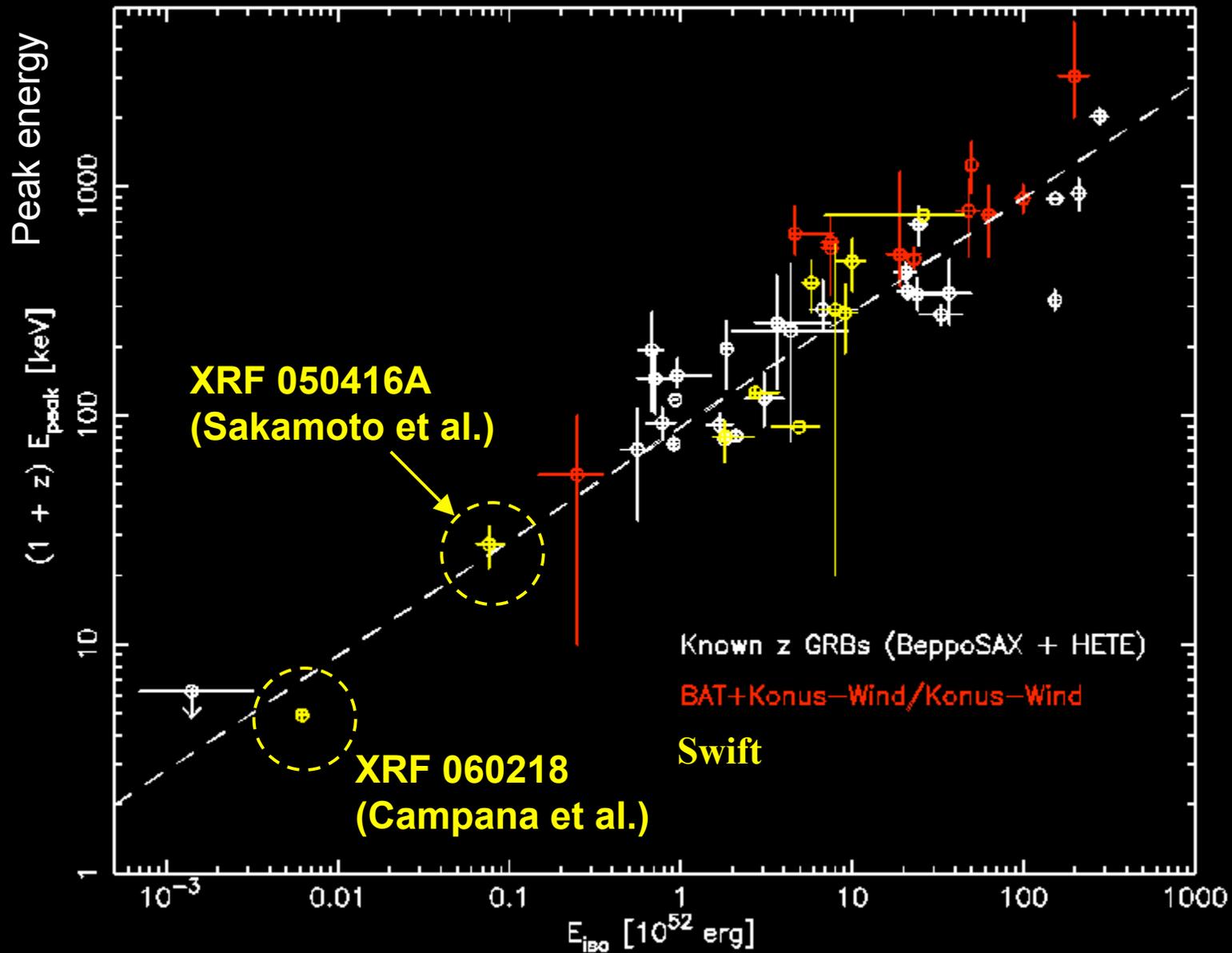
1. IR detection and spectroscopy
2. Deep optical observations
3. Indication of “possible high-z” from the Swift data

**- Luminosity indicators**      Measurement of Dark energy (e.g. Schaefer)

GRB properties ( $E_p$ ,  $V$ ,  $Lag$ , ..) + redshift

1. Variety of redshifts  
(importance of low-z GRBs)
2. Optical/IR/radio detection of AG
3. Spectroscopy

# $E_{\text{peak}} - E_{\text{iso}}$ relation (Swift)



# Summary

## ■ Point Swift to the **anti-Sun direction**

### **Double the number of redshifts**

- XRT: prompt position but no long follow-up
- Q1. How long do we need to cool XRT to get the position?  
(Is cooling XRT during the SAA period enough?)
- Q2. Is it possible to change the threshold of a mode switching?

## ■ Screening “possible” high redshift GRBs from the BAT prompt emission properties

- We can **alert possible high-z ( $z > 4$ ) GRBs** ( $\sim 5$  GRBs year<sup>-1</sup>)

# Sun angle distribution (excluding SAA)

